

GENERATING LOW-TEMPERATURE LAYERS WITH INFRARED HEATING

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Cryogenic targets for the National Ignition Facility require uniform solid layers inside spherical capsules at temperatures ~ 1.5 K below the triple point of hydrogen. Uniform layers have been successfully formed near the triple point. However, upon subsequent cooling the layers degrade. We report here recent attempts to form uniform deuterium hydride (HD) layers 1.5 K below the triple point using infrared (IR) radiation. Pumping the IR collisionally induced vibration-rotation band of solid HD contained inside a transparent plastic shell generates a volumetric heat source in the HD lattice. This in turn allows the formation of a spherical crystalline shell of HD inside the transparent plastic shell. HD layers ~ 50 μm thick have been formed near the triple point and slowly cooled 1.5 K under high IR power without layer degradation.

KEYWORDS: *infrared heating, cryogenics, deuterium hydride*

I. INTRODUCTION

Smooth and uniform 50- to 300- μm -thick deuterium-tritium (D-T) layers on the interior of 1- to 3-mm-diam spherical capsules are required for ignitable inertial confinement fusion targets for the National Ignition Facility¹ (NIF). One way to form these layers is to initially freeze D-T to an anisotropic multicrystalline solid inside a capsule and then allow the tritium decay volumetric heat generation $Q_{\text{D-T}} = 0.05$ W/cm³ to symmetrize the layer. The volumetric heating causes the thicker regions of the solid to have a higher temperature and thus a higher vapor pressure. This results in a redistribution of the solid until the inner D-T surface is isothermal. The time

constant for this redistribution (when no ³He is present) is $\tau_o = H_s \rho / Q_{\text{D-T}} = 23$ min, where H_s is the heat of sublimation (J/mole), and ρ is the density (moles/cm³) (Ref. 2). If bulk heating were the only factor controlling the layer profile, D-T would form a uniform spherical layer inside an isothermal spherical capsule. However, crystal surface stiffness contributes to the final layer shape. The surface structure of a multicrystalline D-T film is a function of the D-T bulk heating rate and the distribution of crystallite sizes and orientations, which are determined by the initial nucleation and growth.³

Several implosion hydrodynamics issues of cryogenic targets can be investigated with deuterium hydride (HD) or deuterium (D₂) without the hazards of handling high-pressure tritium. To generate bulk heat for mass redistribution without tritium radiation, we pumped the collisionally induced infrared (IR) vibration-rotation band. We controlled the solid volumetric heating rate Q_{IR} and thus the redistribution rate $1/\tau_{\text{IR}}$ and surface roughness by controlling the incident IR intensity. The maximum Q_{IR} for hydrogen isotopes is limited by the vibrational relaxation time to $Q_{\text{IR}} \sim 1000 Q_{\text{D-T}}$ (Ref. 4). These large Q_{IR} values enabled us to control the hydrogen layer profile as well as to explore the redistribution process of laser-heated solids.

We have previously reported on experiments where we generated a 4π illumination of HD inside a 1-mm outside diameter (o.d.) spherical capsule and observed the evolution of a uniform solid layer as a function of time and IR intensity.⁵ Initial IR layering experiments formed layers from the solid. However, layers formed this way were extremely rough. In order to improve the quality of the inner solid surface, a second technique was developed. A solid HD sample was first illuminated with IR radiation and then slowly warmed until the solid melted. As soon as the solid melted, the temperature was lowered until solid started reforming in the shell. At that point the sample temperature was held constant until the solidification process was complete. This process was developed using volumetric heating rates on the order of

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eight times the volumetric heating rate of D-T ($8Q_{D-T}$) on layers $150\text{ }\mu\text{m}$ or greater in thickness. Results were shown of a uniform and relatively thick HD layer formed just below the triple point of HD.

In the case of both the tritiated as well as nontritiated hydrogen samples, smooth layers are formed near their respective triple points.⁶ This is, however, not the NIF design point temperature. Attempts to continue the slow cool process in D-T to 1.5 K below the triple point have not been successful. The layers tend to remain smooth to 0.5 to 0.8 K below the triple point at best before starting to degrade. We report here on the results of successful slow cooldown experiments conducted on HD samples illuminated with IR radiation at volumetric heating rates ranging from ~ 1 to $50Q_{D-T}$.

II. EXPERIMENTAL DESIGN

II.A. Experimental Design and Layout

Figure 1 contains a sketch of the experimental layout. For these experiments a $965\text{-}\mu\text{m}$ o.d. \times $891\text{-}\mu\text{m}$ inside diameter (i.d.) deuterated plasma polymer (CD) shell was placed in the center of a 25-mm i.d. integrating sphere. The fill line was composed of two sections: a 5-mm -long tapered glass tube attached to a 7.2-mm -long stainless steel tube. The glass tube was $100\text{-}\mu\text{m}$ o.d. \times $70\text{-}\mu\text{m}$ i.d. with one end drawn down to $\sim 30\text{-}\mu\text{m}$ o.d. to fit into a $30\text{-}\mu\text{m}$ o.d. hole drilled into the shell. The glass

tubing was epoxied into the shell and, after aligning the shell in the integrating sphere viewport, the other end of the tubing was epoxied into the stainless steel tube. The stainless tube extended through the integrating sphere wall. To minimize IR absorption in the fill line, the glass tube was coated with $\sim 1000\text{ }\text{\AA}$ of gold. The laser used for these experiments was a mid-IR optical parameter oscillator (OPO) laser with an output power $>300\text{ mW}$.^a

The integrating sphere was a vacuum-tight enclosure made from aluminum with four window ports and an Infragold coating^b on the inner surface. The Infragold surface is $>96\%$ reflecting at wavelengths of interest. The diffuse reflections off the rough integrating sphere surface reduced laser-induced coherence effects and produced a uniform 4π IR illumination at the capsule. The IR beam was coupled into an optical fiber at the laser output port to transport the IR into the integrating sphere. An SMA bulkhead feedthrough connected the fiber from the laser to the fiber inside the cryostat. An off-axis hole was drilled in the integrating sphere parallel to the capsule fill line but offset by 9.65 mm . One end of a polished optical fiber was positioned with the fiber tip flush to the integrating sphere surface. A leak-tight epoxy seal on the outside of the integrating sphere sealed the fiber in place. Expanding from the fiber tip, the IR beam traversed the integrating sphere parallel to the shell fill line, but the fiber offset prevented the beam from hitting the capsule before it hit and diffusely scattered from the integrating sphere wall. The integrating sphere was thermally and mechanically attached to the cold tip of a helium flow cryostat. To cool the shell, ~ 1.5 Torr of helium exchange gas filled the integrating sphere. The helium pressure was set to a low value to prevent convection from adversely affecting the layer symmetry.

The layer was monitored along a viewing axis perpendicular to the capsule fill line by backlighting the capsule through one port and using a Questar microscope on the opposite port for viewing. Images were acquired periodically during layer formation and subsequent cooling using a Photometrics 1317×1035 array charge-coupled device (CCD) camera attached to the microscope. Real time monitoring of the layer was done with a Cohu camera mounted on the microscope's auxiliary port.

II.B. Data Collection and Analysis Procedures

The first requirement was to relate injected IR laser power to power absorbed by the HD. To determine the bulk heating rate, we measured the redistribution time constant⁷ τ_{IR} . This was accomplished by filling the capsule with liquid HD, rapidly cooling it to $\sim 15\text{ K}$ to

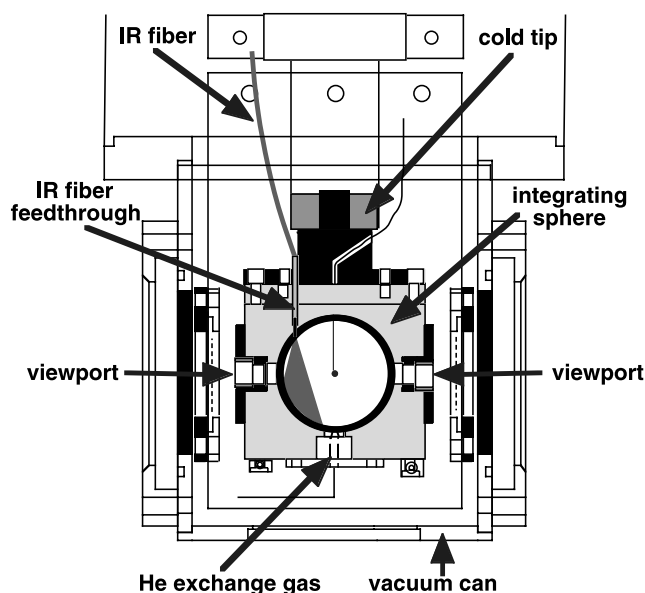


Fig. 1. Experimental layout indicating the viewing axis and the IR fiber. The shaded area inside the integrating sphere represents the spread of the injected IR beam.

^aAculight Corporation, 11805 North Creek Parkway South, Bothell, Washington 98011.

^bInfragold is an electrochemically plated, diffuse, gold-metallic coating from Labsphere, North Sutton, New Hampshire 03260.

freeze the HD, and then turning on the IR and collecting images of the evolving layer with time. Typically, the HD thickness was small near the top of the shell but grew exponentially with time to form a uniform layer. The layer thickness h was measured on each image, the results plotted as a function of time t , and the data fit to an exponential curve of the form $h = a - be^{-t/c}$, where a , b , and c were fit parameters. The exponential time constant is related to the IR volumetric heating rate Q_{IR} by $\tau_{\text{IR}} = H_s \rho / Q_{\text{IR}}$. The IR power was measured by disconnecting the optical fiber at the cryostat feedthrough connector and attaching it to a power meter. For these experiments, ~ 6 mW of laser power available at the cryostat feedthrough resulted in a redistribution time in HD equal to the redistribution time in D-T. In other words, 6 mW corresponded to a volumetric heating rate of $1Q_{\text{D-T}}$. By adjusting the OPO laser's output power, we could generate volumetric heating rates in the range of ~ 1 to $50Q_{\text{D-T}}$.

The typical layering process involved injecting IR into the integrating sphere, raising the temperature of the capsule until the solid just finished melting, and then slowly lowering the temperature in ~ 0.005 K steps until freezing onset occurred, at which point the temperature was no longer lowered. If the solidification process appeared to stop, then the temperature was lowered in 0.005 K steps until the HD completely solidified. Typically, the solidification process finished before the temperature was lowered 0.1 K below the melting point. Once the HD solidified, the temperature was lowered 1.5 K at a constant cooling rate in steps of either 0.020 K every 3 min (20 mK/3 min) or 0.010 K every 3 min (10 mK/3 min). While the temperature was lowered, images of the layer were captured and stored every 6 min.

Roughness of the inner solid surface was determined by performing a Fast Fourier Transform (FFT) analysis of the brightband in the stored CCD camera images. The brightband was produced by total internal reflection of light rays off the inner solid surface.^c Software has been developed that locates the brightband in the image.⁸ The resultant one-dimensional estimates of the solid thickness variations were Fourier transformed giving values for the power spectrum and root-mean-square (rms) of perturbations on the solid layer. Early surface roughness calculations located the position of the inner edge of the brightband to determine these values. Computer raytrace modeling has since shown that a better estimate of the power spectrum and rms value is made by locating the position of the center of the brightband.⁹ The data presented in this paper use the center of the brightband position relative to the outer edge of the

capsule to fit Fourier modes 1 through 128. Due to the perturbation caused by the fill tube, the defect in the layer near the fill tube ($\sim 17\%$ of the brightband) was excluded before the brightband analysis was performed.

III. RESULTS

The goal of these experiments was to determine if a uniform HD layer formed and maintained under constant IR illumination near the triple point of HD could be kept from degrading while slowly lowering the temperature 1.5 K. The determination of whether or not the layer degraded was done by a combination of both qualitative and quantitative analyses of the images. The quantitative information from our images was acquired by FFT analysis of the brightband. If the layer starts to degrade, then the rms value for that layer will increase. Since there is currently no quantitative analysis performed on the center of the images, we looked qualitatively for the growth of features as the layer cooled. Changes in features were regarded as the growth of defects and interpreted as layer degradation.

Figure 2 contains a plot of the surface roughness results from one of the slow cooldown experiments. The laser power was set to generate a volumetric heating rate of $10Q_{\text{D-T}}$, and the temperature was lowered in 20 mK/3 min steps. The plot shows the rms values from the brightband FFT analysis of the images for modes 4 to 100 plotted versus temperature. If the surface roughness did not change with temperature, the data points would fit along a flat line, as indicated by the solid line in the

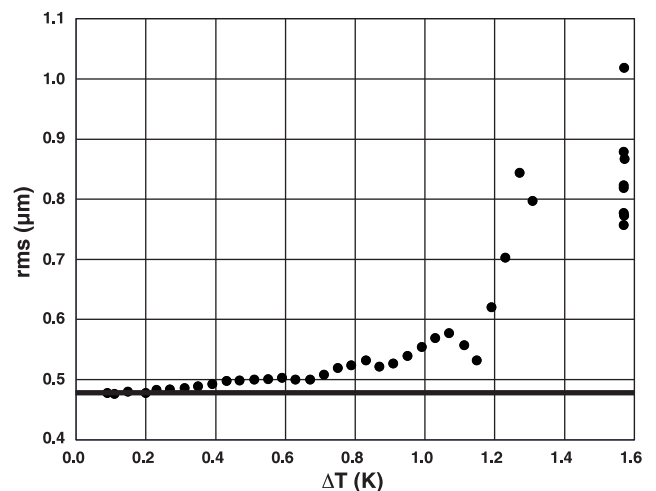


Fig. 2. An example plot of the brightband analysis results for one of the 20 mK/3 min cooling rate experiments. The volumetric heating rate was $10Q_{\text{D-T}}$. The rms value was calculated for modes 4 to 100. The solid line indicates the rms value for the layer if it does not degrade.

^cThe brightband is formed by total internal reflection off the inner solid surface, and its position is determined by the refractive effects of the CD shell as well as the solid HD. The correlation of the inner ice surface with the brightband is discussed further in Refs. 5 and 12.

plot. As one can see in this data, the surface rms value slowly increases $\sim 10\%$ by the time the temperature is lowered 0.8 K. Below this temperature the rms value starts to increase more rapidly, and between 1.3 and 1.55 K the brightband is too discontinuous for analysis with our software. After dropping by 1.57 K, the temperature was no longer lowered, but images were acquired for an additional 43 min. During this time, some annealing of the surface occurred, and we were able to analyze the images acquired during the final 28 min. This annealing was typical of most of the experiments. However, there was never an indication that sufficient annealing would occur with time to decrease the rms value to the value measured after initial layer formation.

Figure 3 shows four of the images acquired during the cooldown. The first image was captured after initial layer formation ($\Delta T = T_{\text{melt}} - T = 0.09$ K). This is our

reference image. During the subsequent cooldown we looked for changes relative to the reference image indicating that the layer was degrading. The next image ($\Delta T = 1.03$ K) does not show any break in the brightband, but features are starting to grow near the bottom of the capsule. The third image ($\Delta T = 1.55$ K), acquired at the next-to-the-last 20 mK temperature step, shows more features near the bottom of the capsule and multiple discontinuities in the brightband. The discontinuities here were too great for analysis. The final image ($\Delta T = 1.57$ K) was acquired after the layer sat at constant temperature for 43 min. One can see that some annealing of the layer has occurred, and one of the discontinuities near the bottom of the capsule has disappeared. However, this last image still shows a layer qualitatively as well as quantitatively rougher than the image at the start of the cooling sequence.

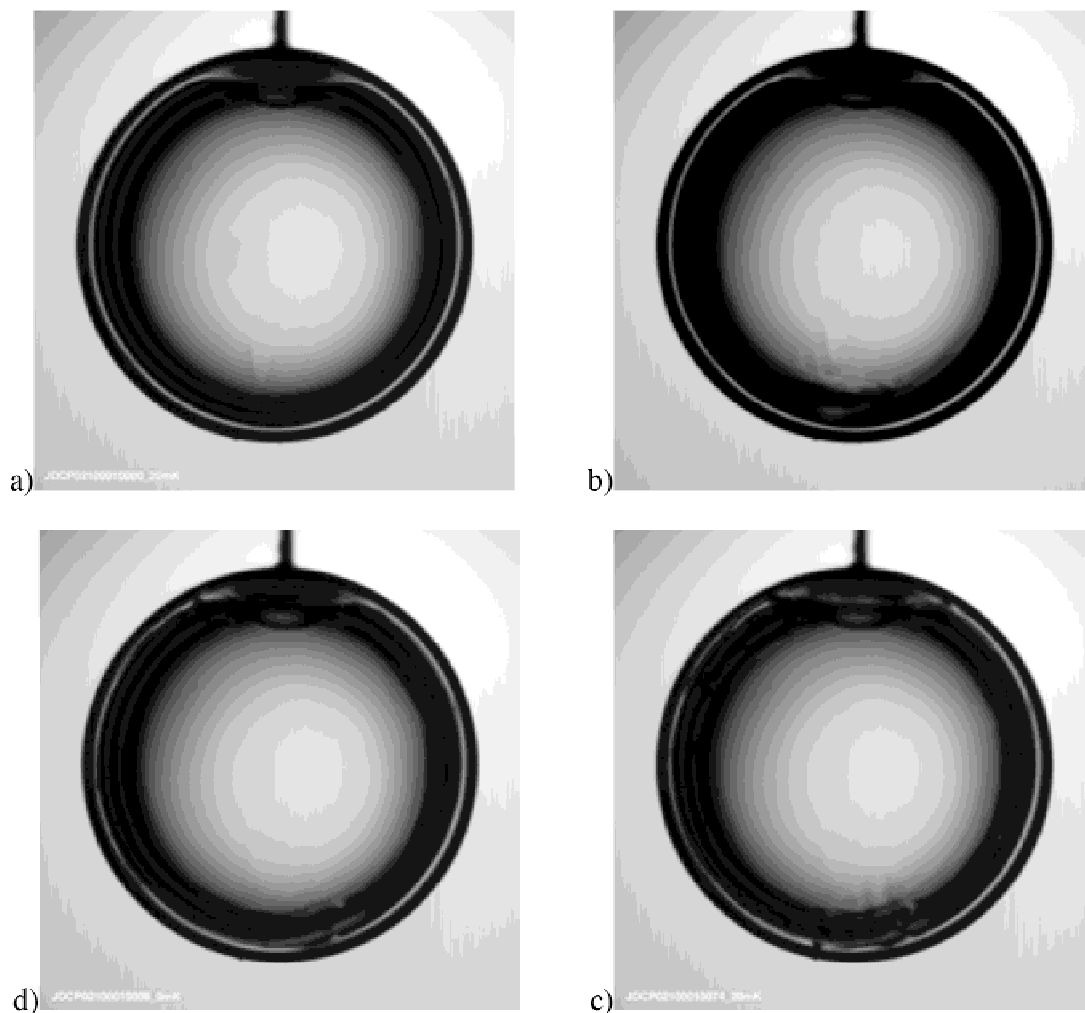


Fig. 3. Corresponding images from the slow cooldown of the layer analyzed in Fig. 2; (a) the layer near the triple point ($\Delta T = 0.09$ K); (b) the layer at $\Delta T = 1.03$ K where additional small features are evident inside the brightband; (c) breaks in the brightband are seen near the end of the cooling cycle ($\Delta T = 1.55$ K); (d) some annealing of the brightband after sitting at 1.57 K for 43 min.

At a cooling rate of 20 mK/3 min, we were not able to maintain the layer quality. We were able to minimize the degree of layer degradation by increasing the volumetric heating rate. Figure 4 contains a pair of images from both a low, $\sim 1.3Q_{D-T}$, and a high, $\sim 46Q_{D-T}$, volumetric heating rate. The layer cooled using the low heating rate shows significant layer degradation. In fact, the layer rms value increased rapidly as the layer was cooled. The rms value (modes 4 to 100) started at $0.6 \mu\text{m}$, but the layer became too rough to analyze by the time the temperature was lowered 0.2 K. On the other hand, the layer cooled using a high volumetric heating rate showed a much slower layer degradation rate, and the rms value ($0.47 \mu\text{m}$ for modes 4 to 100) did not significantly change until near the end of the cooling cycle when the layer had cooled over 1 K.

Based on the results of these measurements, a second set of slow cooldown experiments was conducted at a cooling rate of 10 mK/3 min. At this cooling rate it now took 7.5 h to lower the temperature 1.5 K. Using this slower cooling rate with a relatively high volumetric heating rate we were able to maintain the layer quality through the entire cooldown process. Figure 5 shows the images of a layer taken at the start and the end of a cooldown where the volumetric heating rate was $\sim 34Q_{D-T}$. As one can see, there are no qualitative differences in the two images. Quantitatively, there is no significant difference between the rms values of these two images. Figure 6 shows the power spectra for these two images. Only small variations appear between the mode amplitudes. The table insert lists the rms values for the first several mode ranges. Most of the power is in

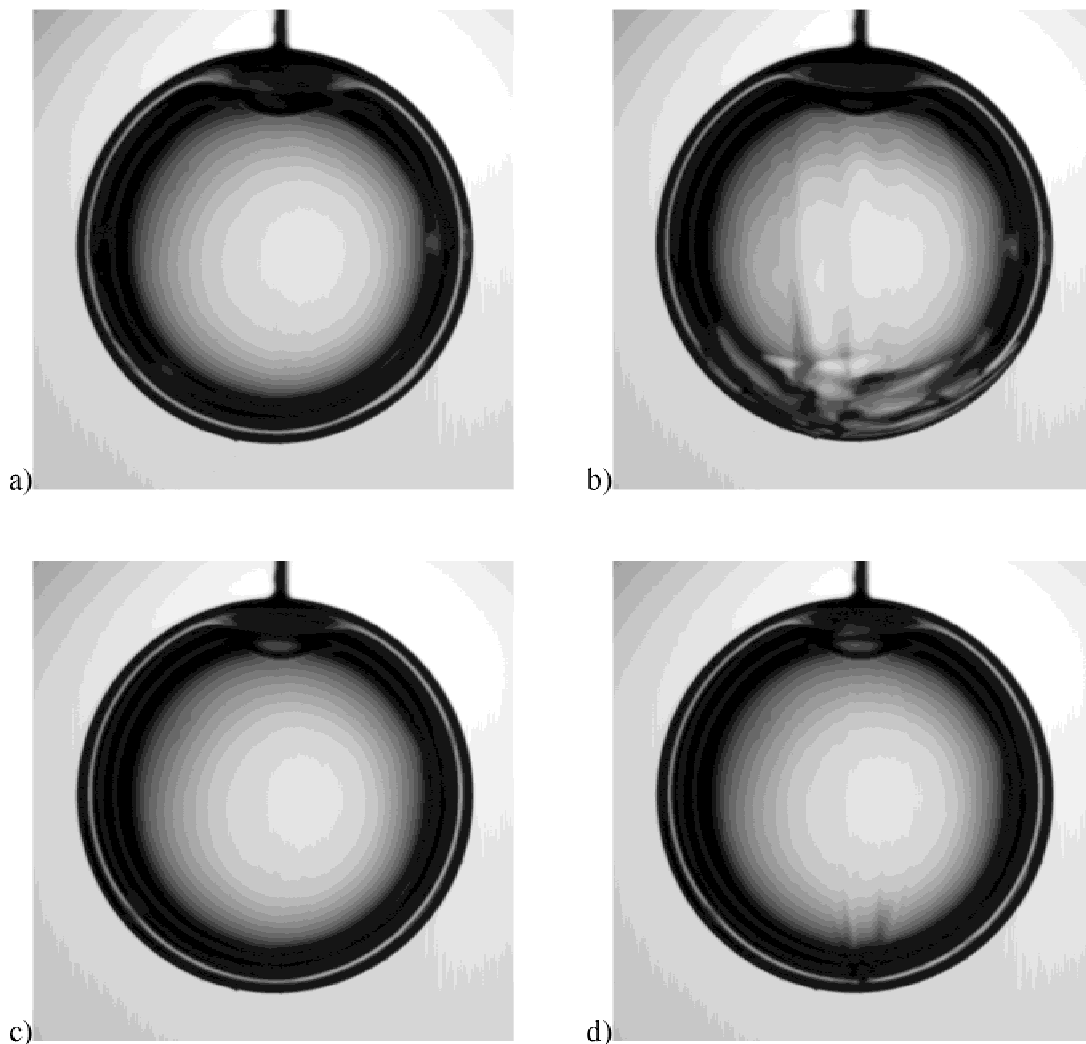


Fig. 4. Images comparing layers formed and cooled at 20 mK/3 min under low (a and b) and high (c and d) IR power; (a) image of a layer formed with a volumetric heating rate of $\sim 1.3Q_{D-T}$; (b) the same layer after cooling 1.5 K. In contrast, (c) is an image of a layer formed at $\sim 46Q_{D-T}$, and (d) is the same layer after cooling 1.5 K.

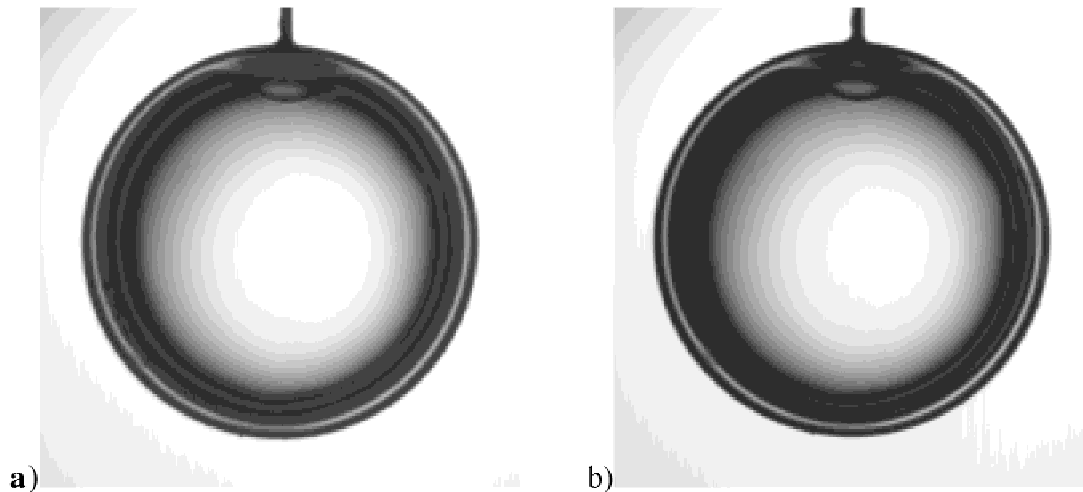


Fig. 5. Images of a layer cooled at 10 mK/3 min under a volumetric heating rate of $\sim 34Q_{D-T}$; (a) the layer just below the triple point; (b) the layer 1.5 K below the triple point.

the first three modes. The difference in rms values between the high ($\Delta T = 0.1$ K) and low ($\Delta T = 1.5$ K) temperature modes is $< 11\%$. For this second set of slow cooldown experiments, the lowest volumetric heating rate at which the layer quality was maintained was $\sim 25Q_{D-T}$.

IV. CONCLUSIONS

We have demonstrated that it is possible to maintain the integrity of a uniform HD layer at 1.5 K below the

triple point by slowly cooling down the layer under IR illumination. The best layers were formed under high IR power, corresponding to heating rates of at least $25Q_{D-T}$, with a cooling rate of 10 mK/3 min. Layers formed and cooled under low IR power and layers cooled at a rate of 20 mK/3 min degraded before the temperature was lowered 1.5 K.

These were a series of experiments to determine the feasibility of the slow-cooling concept. There should be a determination of the minimum volumetric heating rate and maximum cooling rate necessary to maintain the layer to 1.5 K below the triple point. Only two constant cooling rates were used. At a cooling rate of 10 mK/3 min, the cooling cycle took nearly 8 h to lower the layer temperature 1.5 K. It may be possible to cool the layer at a faster rate at the beginning of the cooling cycle and then cool at a slower rate near the end of the cooling cycle. Further parametric studies are needed to determine an optimal cooling protocol.

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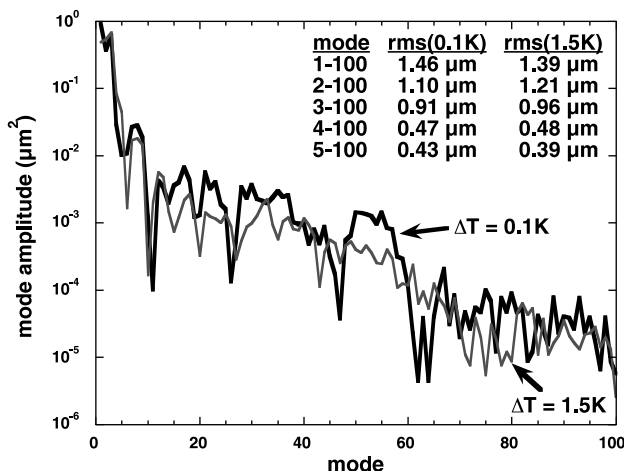


Fig. 6. Power spectra for the two images shown in Fig. 5. The thick line is the power spectrum of the layer after formation ($\Delta T = 0.1$ K) and the thin line is the power spectrum of the layer after cooling 1.5 K. The table insert lists the various mode range rms values for comparison.

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